

# Lowering the noise in a CCD readout system

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# Problem statement and chronology

- CCD readout noise is a limitation to achieve better physics/ astronomy in telescopes or CCD based detectors.
- Can this noise be lowered beyond the levels achieved by the traditional Correlated Double Sampling technique? (i.e.  $\sim 2e^-$ )
- Approach:
  - Take N samples of each pixel at sampling frequency  $f_s$ .
  - Digital Signal Process the data.
- We sampled CCD video data (few lines of an image) using an existing hardware (ESECON) developed in our ESE Dept in CD.
  - Data analysis is still in progress. Some preliminary results.

# CCD noise and CCD readout noise

- CCD noise
  - Dark current is negligible at cryogenic temperatures.
  - Mainly shot noise from signal (Poisson).
- CCD readout noise:
  - $1/f$  noise.
  - Thermal (shot) noise (WGN), limited by high bandwidth amplifiers.
  - Analog to digital conversion noise (quantization noise).
- CCD readout noise sources
  - CCD internal electronics.
  - Readout amplifier chain.
  - Analog to digital converter and associated electronics.

# CCD readout system



- CCD Picture frame adapter. It has a JFET amplifier.
- Flex or normal cable in the Dewar. Opportunities for crosstalk between higher voltage digital signals and clocks to the sensitive CCD video outputs.
- Preamplifier outside the Dewar.
- Digitizer and signal processing board.

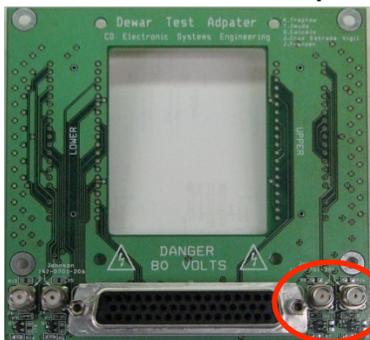
Digitizer (ESECON)



Preamplifier



CCD frame adapter



Cable inside the Dewar

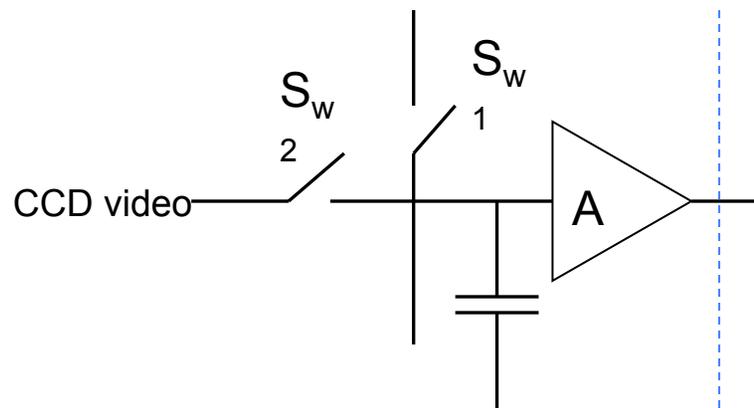


Video output and JFET

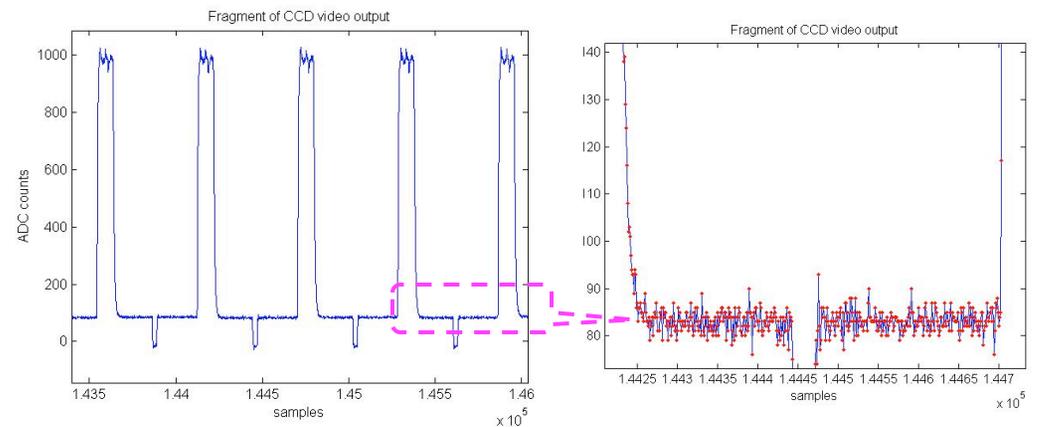
# CCD readout

- CCD pixels are readout by transferring the pixel charge serially to one of the CCD's video outputs.
- The CCD uses timing signals (clocks) to do the charge transfer.
  - Timing signals can be of arbitrary length.
  - Unfortunately, the clocks show up in the video data stream.
- A pixel readout is done in two steps:
  - 1<sup>st</sup> the CCD video output capacitor is reset (CCD internal). This sets a pedestal value.
  - 2<sup>nd</sup> the pixel charge is transferred to the same capacitor.
- The actual pixel signal is calculated as a function of the difference between the pedestal charge and the pixel charge.

Simplified diagram of CCD video output



CCD video output with clocks



# Most commonly used CCD readout system

- Correlated Double Sampling (CDS) performs

$$s(t) = \frac{1}{T} \int_0^T [x(t) - p(t)] dt \quad \Rightarrow \quad s(t) = \frac{1}{T} \left[ \int_0^T x(t) dt - \int_{t_0}^{T+t_0} p(t) dt \right] \quad \text{for an integration window } T$$

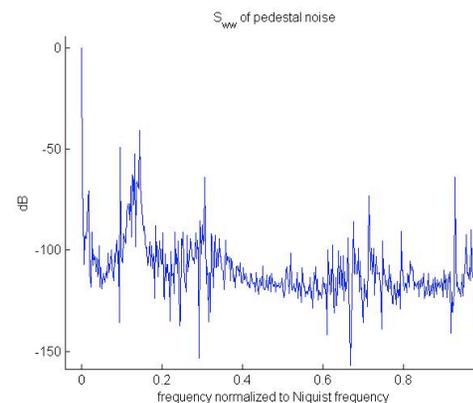
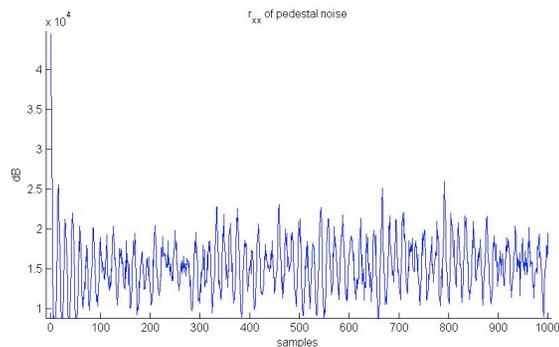
Where  $x(t)$  is the pixel and  $p(t)$  is the pedestal.

Let:  $x(t) = x_0 + n_x(t)$  and  $p(t) = p_0 + n_p(t)$        $n_x$  and  $n_p$ : signal and pedestal noise

If the noise is white and Gaussian  $\sim N(0, \sigma^2)$  the CDS is the optimum estimator.

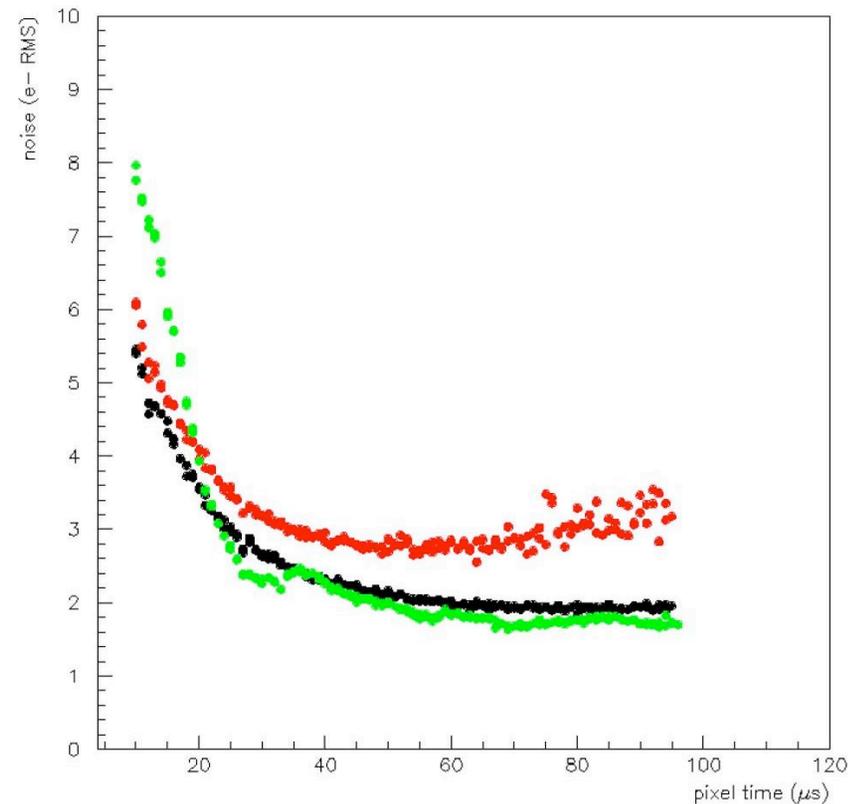
$$\hat{s}(t) = x_0 - p_0 \Big|_{T \rightarrow \infty} \quad \text{or} \quad \sigma_{\hat{\theta}_s}^2 \rightarrow 0 \Big|_{T \rightarrow \infty}$$

Unfortunately there is correlated noise in the low frequency spectrum. The noise does not integrate to 0 when T is made larger.



# CCD noise measured with the Monsoon system

- The noise increases for long pixel observations  $T > 100 \mu\text{s}$ .
- Besides it would be advantageous to have a low noise readout system that reads fast too.



# The classical CDS

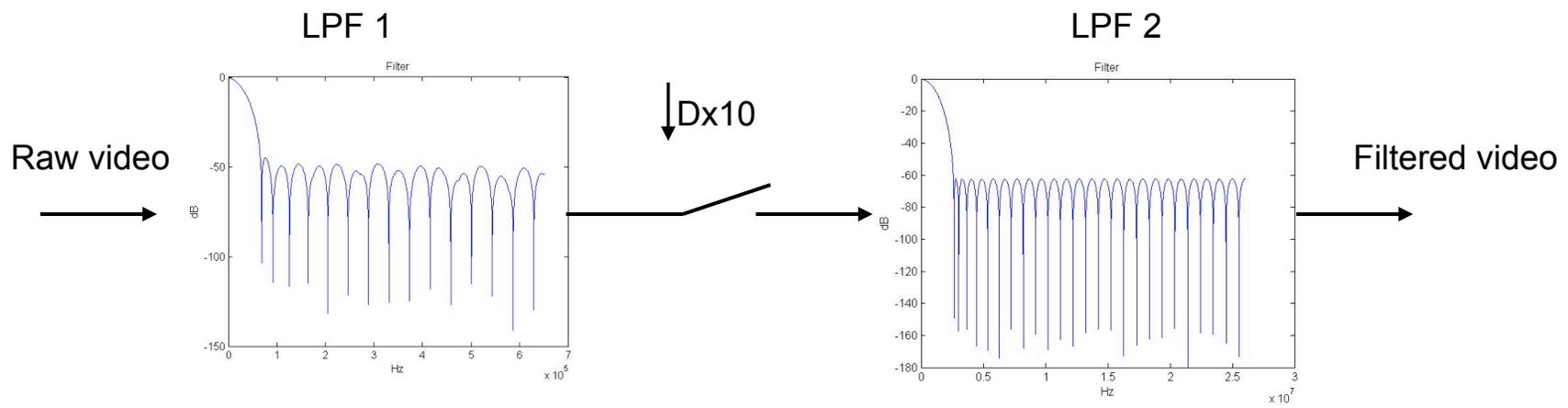
- **Advantages:**
  - Generates a single data point per pixel.
  - The pixel estimation gets rid of the pedestal value which is small but still a random number.
  - The CDS filtering is implemented by a simple analog circuit.
- **Disadvantage:**
  - CDS integrates the entire noise spectrum  $[0, \infty)$  over a limited time  $T$ . Low frequency noise does not integrate to 0. The estimator variance does not go to zero as  $T \rightarrow \infty$ .

# Alternative to the CDS CCD Readout

- Sample the video output.
  - Take  $N$  samples of the pedestal and the pixel value during the observation time  $T$ .
  - Digital Signal Process the data.
  - Apply a digital CDS to the processed data.
- Digital signal processing methods used:
  - Finite Impulse Response filtering.
  - Estimation using a linear model.
  - The pixel estimation problem is the classical problem of estimating a DC value in noise.

# FIR filtering

- Two stage FIR filter with decimation.
  - Used to filter the high frequency noise content.
  - The data is oversampled and later decimated as a way to reduce the quantization noise of the analog to digital converter.
  - Lowers the noise of 97% of the spectrum (Nyquist domain) by at least 40dB (a factor of 100).
  - The FIR filters improve the estimation by only 30 to 40% with respect to CDS.

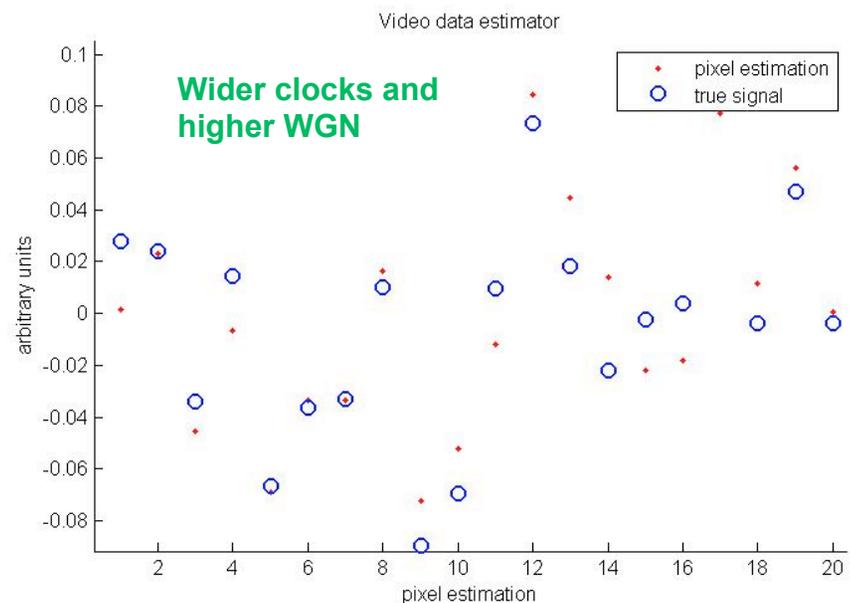
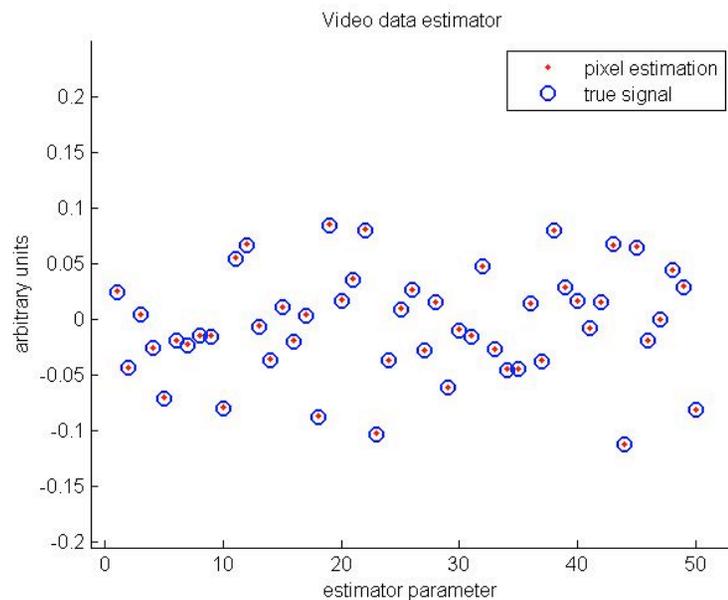


# Estimators

- So far we used two estimators:
  - One estimates the low frequency noise content in the time domain and the second one in the frequency domain.
  - The estimation problem is linear and the inversion of the estimation matrix is done only once and a priori.
  - Both estimators were tested on simulated data. The time domain estimator was tested on a small set of real CCD sampled data.

# Simulated data

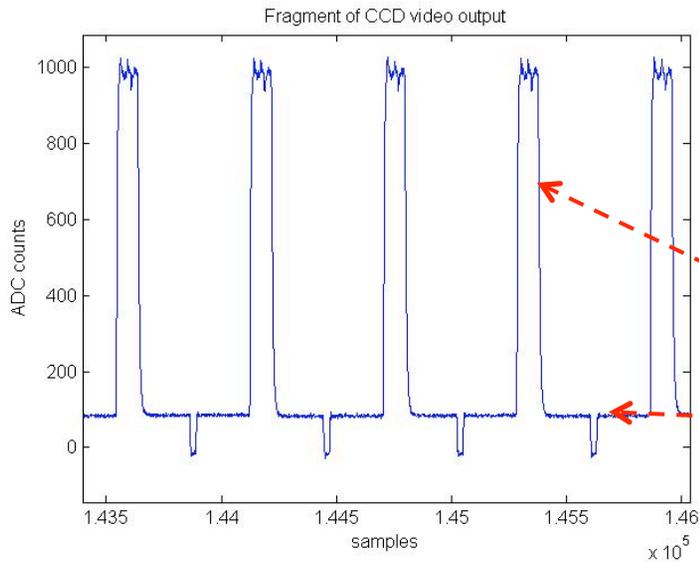
- Random pixel data with additive noise.
- The noise follows an  $1/f$  power spectrum with 500 to 1000 frequency components.
  - Amplitude and phases are randomized.
  - These are the noise components that we estimate.
- Small amplitude WGN noise, to test the robustness of the solution.
  - This is the high frequency noise that we are not estimating.
- We also include large the noisy clocks.



# Observations

- The clocks are removed from the sampled data by skipping the points in time domain.
  - Clock gaps create numerical problems in the estimation process.
  - Clock gaps can become very large unless the pixel observation time is made longer.
- Some WGN is tolerated and the error increases linearly with the square root of the noise power.
- We must estimate all the frequencies in the filter's bandpass.

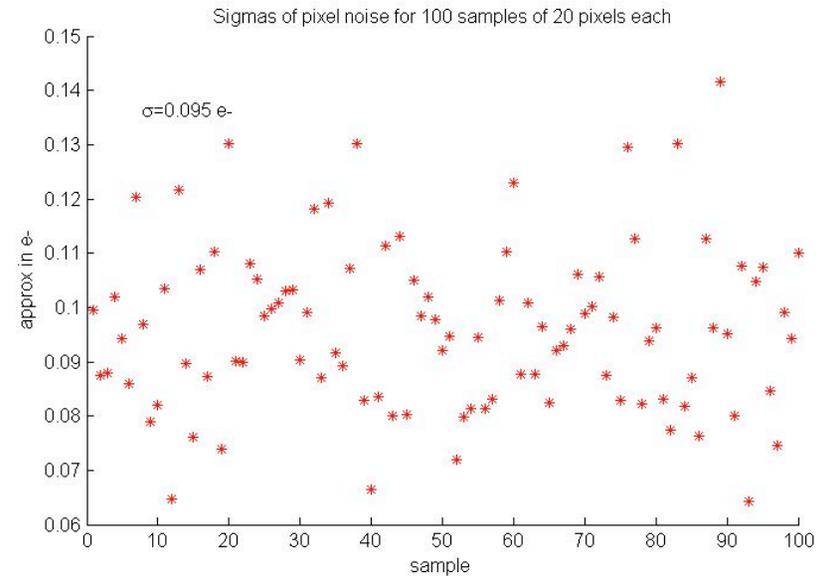
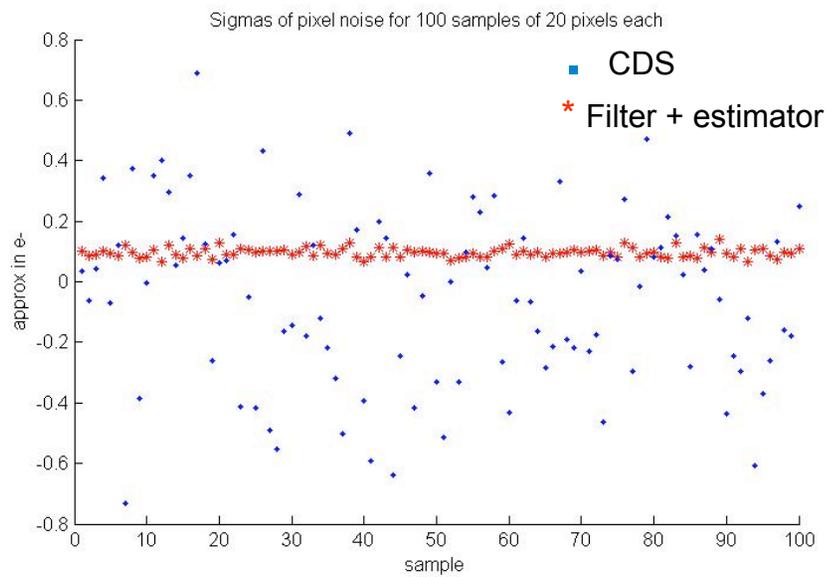
# Real CCD data



- What we see in simulation is verified.
- Clocks are huge compared to the signal levels that we are trying to measure.
  - Clocks  $\sim 1000$ .
  - Signals  $\sim 0.01$  in this arbitrary scale.
- Clocks gaps are too large.

# Real CCD data

- Plots below show the sigma of the estimation error for and a comparison with the CDS (left).
- This is our current best results that could not be reproduced consistently in all our data.



# Future work

- Make the filter-estimator work consistently.
- Process an entire image and see if it works for real.
- Hardware optimization.
  - Several noise reduction techniques.
  - I do not think that we need a new sampling card yet.